

DAMTP-2017-7, ICCUB-17-04

Tachyons in the Galilean limit

Carles Batlle¹, Joaquim Gomis²,
Luca Mezincescu³ and Paul K. Townsend⁴

¹Departament de Matemàtiques and IOC, Universitat Politècnica de Catalunya
EPSEVG, Av. V. Balaguer 1, E-08808 Vilanova i la Geltrú, Spain

email: carles.batlle@upc.edu

²Departament de Física Quàntica i Astrofísica
and Institut de Ciències del Cosmos (ICCUB), Universitat de Barcelona,
Martí i Franquès 1, E-08028 Barcelona, Spain

email: gomis@fqa.ub.edu

³Department of Physics, University of Miami,
P.O. Box 248046, Coral Gables,
FL 33124, USA

email: mezincescu@physics.miami.edu

⁴Department of Applied Mathematics and Theoretical Physics,
Centre for Mathematical Sciences, University of Cambridge,
Wilberforce Road, Cambridge, CB3 0WA, U.K.

email: P.K.Townsend@damtp.cam.ac.uk

ABSTRACT

The Souriau massless Galilean particle of “colour” k and spin s is shown to be the Galilean limit of the Souriau tachyon of mass $m = ik$ and spin s . We compare and contrast this result with the Galilean limit of the Nambu-Goto string and Green-Schwarz superstring.

1 Introduction

The action describing a free non-relativistic point particle is Galilean invariant with the particle's mass appearing as a central charge in the Poisson bracket algebra of the corresponding Noether charges; this enlarged algebra is called the Bargmann algebra. A strictly Galilean dynamical system that realizes the Galilei algebra without this central charge is a “massless Galilean” system; the concept and terminology are due to Souriau [1], who also provided a simple example, which has applications to spinoptics [2,3]: the massless Galilean particle of “colour” k and spin s .

It was recently shown that the Nambu-Goto string admits a strictly Galilean limit, and the same is true for any Dirac-type p -brane for $p > 0$, so these provide further examples of massless Galilean systems [4]. It has also been shown that the Green-Schwarz superstring admits a super-Galilean limit in which the Galilei algebra is enlarged to a superalgebra [5]. Although this superalgebra does have a central charge, the Galilei subalgebra does not, so the Galilean superstring provides an example of a “massless super-Galilean” system.

Curiously, the limiting procedure that leads to the Galilean p -brane does not apply for $p = 0$. The Galilean massless particle is *not* the Galilean p -brane for $p = 0$; in other words, it is not a Galilean limit of the massive relativistic particle. Could it be a limit of the *massless* relativistic particle? In paragraph 14.54 of the English edition of his book, Souriau affirms that it is, but he says that the limit is “of a different kind” that “gives rise to a family of distinct non-relativistic particles, each one labeled by a color k ” [1]. Unfortunately, Souriau does not give details, and this is also true of a statement of relevance here that he makes in the very next paragraph of his book: “As for tachyons, it does not seem that one can obtain a non-relativistic limit for them”.

In this paper we show that the massless Galilean particle of colour k is a limit of the relativistic *tachyon* of imaginary mass $m = ik/c$ (where c is the speed of light). This explains why the massless Galilean particle action cannot be obtained by choosing $p = 0$ in the Galilean p -brane action of [4]: the relativistic starting point for the former is not the $p = 0$ case of the relativistic starting point of the latter. The massless Galilean particle and the Galilean p -brane for $p > 0$ are two quite different massless Galilean systems. Further differences become apparent when one considers the extension to massless super-Galilean systems, but we postpone this discussion to the end of the paper.

We begin with a brief review of the massless Galilean particle in a notation that is convenient for our purposes, comparing and contrasting it with the Galilean string. We then review the massive relativistic particle with spin incorporated via the manifestly Lorentz invariant “Souriau 2-form”, before discussing its tachyonic version and taking the Galilean limit to recover the massless Galilean particle. We comment on the incorporation of spacetime supersymmetry in our concluding discussion.

2 Galilean massless particle

The phase space of the massless Galilean particle is parametrized by position 3-vector \mathbf{x} and time scalar t , and their conjugate momenta \mathbf{p} and E , subject to one phase-space constraint. The phase-space action is¹

$$S = \int d\tau \left\{ \mathbf{p} \cdot \dot{\mathbf{x}} - Et - \frac{1}{2}e(|\mathbf{p}|^2 - k^2) \right\} - sS_{WZ}, \quad (2.1)$$

where the overdot indicates a derivative with respect to the arbitrary worldline parameter τ , and S_{WZ} is what we now customarily call a Wess-Zumino action; it is derived from the phase-space 2-form (the exterior product of forms is implicit)

$$\Omega_{WZ} = \frac{1}{2k^3} \mathbf{p} \cdot d\mathbf{p} \times d\mathbf{p}. \quad (2.2)$$

This 2-form is closed ($d\Omega_{WZ} = 0$) as a consequence of the phase-space constraint imposed by a Lagrange multiplier e :

$$|\mathbf{p}|^2 = k^2. \quad (2.3)$$

Souriau did not write down this action as he preferred to work directly with the symplectic 2-form

$$\Omega = d\mathbf{p} \cdot d\mathbf{x} - dE dt - \frac{s}{2k^3} \mathbf{p} \cdot d\mathbf{p} \times d\mathbf{p}. \quad (2.4)$$

Inversion of Ω on the constraint surface yields the canonical Poisson brackets

$$\{E, t\}_{PB} = 1, \quad \{x^i, p_j\}_{PB} = \delta_j^i, \quad \{x^i, x^j\}_{PB} = -\frac{1}{k^3} s \varepsilon^{ijk} p_k, \quad (2.5)$$

where $\{x^i, p_i; i = 1, 2, 3\}$ are the cartesian components of \mathbf{x} and \mathbf{p} . Notice that these relations imply, for non-zero s , that the space coordinates become non-commuting operators in the quantum theory.

The manifest Galilean invariance of Ω shows that the action (2.1) is Galilean invariant up to a surface term, despite the fact that there is no manifestly Galilean invariant expression for S_{WZ} . The corresponding Noether charges are

$$H = E, \quad \mathbf{P} = \mathbf{p}, \quad \mathbf{G} = \mathbf{p}t, \quad \mathbf{J} = \mathbf{x} \times \mathbf{p} + \frac{s}{k} \mathbf{p}. \quad (2.6)$$

A simple way to verify the expression for \mathbf{J} is to consider the variation of Ω induced by an infinitesimal rotation with parameter $\boldsymbol{\omega}$ (i.e. $\delta\mathbf{x} = \boldsymbol{\omega} \times \mathbf{x}$ etc.). One finds that

$$\delta\Omega = d(d\boldsymbol{\omega} \cdot \mathbf{J}), \quad (2.7)$$

¹The constraint is that of the massless relativistic particle with $p_0 = k$, which is the classical analog of the restriction of solutions of the wave equation to those of frequency k , hence the “colour” terminology for k .

where \mathbf{J} is as given. This confirms² that s represents spin³. We remark that the same WZ term, but with \mathbf{p} replaced by an $SO(1, 2)$ vector, was used in [6] to incorporate spin in the action for a *relativistic* particle in a 3-dimensional *Minkowski* spacetime.

Using the Poisson bracket relations (2.5), it may be verified that the Noether charges (2.6) span the Galilei algebra. In particular, one finds that

$$\{P_i, G_j\}_{PB} = 0, \quad (2.8)$$

which implies that the total momentum is boost-invariant! Compare this state of affairs with that of the standard non-relativistic point particle: its mass m appears as a central charge in this Poisson bracket relation, implying that the total momentum is not boost invariant, as one would expect. The absence of this central charge is the characteristic feature of a massless Galilean system.

2.1 Comparison with the Galilean string

Let us pause to make a comparison (for zero spin) with the (closed) Galilean string [4]. In this case all canonical variables are periodic functions of the string coordinate σ , and the Galilean Noether charges are

$$H = \oint d\sigma E, \quad \mathbf{P} = \oint d\sigma \mathbf{p}, \quad \mathbf{G} = \oint d\sigma \mathbf{p} t, \quad \mathbf{J} = \oint d\sigma \mathbf{x} \times \mathbf{p}. \quad (2.9)$$

The phase space constraint is found from a Galilean limit of the string mass-shell constraint $p^2 + (Tx')^2 = 0$, where the prime indicates a derivative with respect to σ , and this limit yields

$$|\mathbf{p}|^2 = (Tt')^2. \quad (2.10)$$

If the string is wound n times around the “time direction”, thus allowing the gauge choice $t' = n$, one can show that the total momentum \mathbf{P} satisfies the bound [5]

$$|\mathbf{P}|^2 \leq n^2. \quad (2.11)$$

This has non-trivial solutions if $n \neq 0$, but there is no particle analog of this possibility. The mass-shell constraint for a particle of mass m is $p^2 + (mc)^2 = 0$ (we use the “mostly plus” Minkowski metric signature) and the same limit yields

$$|\mathbf{p}|^2 = -(mc)^2, \quad (2.12)$$

which has no solutions for real non-zero m (and only the trivial solution $\mathbf{p} = \mathbf{0}$ for $m = 0$). However, it does have solutions if we allow m to be imaginary, in which case the relativistic particle is a tachyon.

We shall now pursue this idea for a relativistic progenitor of the massless Galilean particle in the context of a classical description, again due to Souriau [1], of a *relativistic* particle of mass m and spin s in a 4-dimensional Minkowski background.

²The variation of the Lagrangian 1-form is a total derivative for constant ω , and we read off the corresponding Noether charge from the derivative of ω .

³What we are calling spin is decomposed by Souriau into a magnitude that he calls spin and a sign that he calls “helicity”.

3 The Souriau spinning particle

Souriau's Lorentz covariant description of the massive spinning particle requires the introduction of an independent “polarization” 4-vector (w) in addition to the particle's position 4-vector (x) and momentum 4-vector (p). These are subject to the three constraints

$$p^2 = -(mc)^2, \quad p \cdot w = 0, \quad w^2 = (mc)^2 s^2. \quad (3.1)$$

Now we introduce the 2-form

$$\Omega = dp^\mu dx_\mu + \Omega_S, \quad (3.2)$$

where the second term is the “Souriau 2-form”

$$\Omega_S = \frac{1}{2p^2} \varepsilon^{\mu\nu\rho\sigma} w_\rho p_\sigma \left(\frac{1}{p^2} dp_\mu dp_\nu + \frac{1}{w^2} dw_\mu dw_\nu \right). \quad (3.3)$$

This 2-form is closed (and hence so is Ω) as a consequence of the constraints, as follows from the following lemma:

- **Lemma:** Given two 4-vectors (u, v), the 2-form

$$\omega = \frac{1}{2} \varepsilon^{\mu\nu\rho\sigma} v_\rho u_\sigma (du_\mu dv_\nu - dv_\mu du_\nu) \quad (3.4)$$

is closed if

$$u^2 = -1, \quad v^2 = 1, \quad u \cdot v = 0. \quad (3.5)$$

To prove this lemma, we first observe that these constraints imply

$$u \cdot du = 0, \quad v \cdot dv = 0, \quad u \cdot dv + v \cdot du = 0. \quad (3.6)$$

We now choose a Lorentz frame for which

$$u_\mu| = (1; 0, 0, 0), \quad v_\mu| = (0; 0, 0, 1), \quad (3.7)$$

where the $|$ notation indicates that this choice is made at one point; i.e. it is not assumed to hold for du and dv . However, the derived constraints (3.6) in this frame are

$$du_0 = 0, \quad dv_3 = 0, \quad dv_0 = du_3. \quad (3.8)$$

Using both (3.7) and (3.8), a straightforward calculation yields $d\omega = 0$.

By observing that $\Omega_S = s\omega$ for $(p, w) = mc(u, sv)$, we conclude that $d\Omega_S = 0$.

It is important to appreciate that Ω is not a “symplectic” 2-form for the 12-dimensional space parametrized by the components of the three 4-vectors (x, p, w) . This is because it is not invertible on this space; it is block diagonal in the basis $\{dx, dp, dw\}$ but the 4×4 (dw, dw) block has p and w as two zero-eigenvalue eigenvectors. However, within the 4-dimensional Minkowski subspace of fixed (x, p) the two

w -dependent constraints determine a (p -dependent) 2-sphere whose tangent vectors are orthogonal to both p and w . To see this it suffices to choose the frame for which $\mathbf{p} = \mathbf{0}$; then $w = (0, \mathbf{w})$ with $|\mathbf{w}|^2 = (mcs)^2$. The pull-back of Ω to the 10-dimensional submanifold of topology $\mathbb{R}^8 \times S^2$ defined by the w -dependent constraints *is* invertible. Its inversion yields a set of canonical Poisson brackets for this phase space, with respect to which the remaining w -independent constraint is first-class, so the physical phase-space is 8-dimensional; in fact it is topologically $\mathbb{R}^6 \times S^2$, where the first factor is the phase space for a free particle in the Euclidean 3-space and the second factor is the spin phase space (as becomes manifest in a bi-twistor formulation [7, 8]).

Finally, to see why the parameter s is the particle's spin, we observe that the infinitesimal Lorentz transformations

$$\delta x^\mu = \Lambda^\mu{}_\nu x^\nu, \quad \delta p_\mu = \Lambda_\mu{}^\nu p_\nu, \quad \delta w^\mu = \Lambda^\mu{}_\nu w^\nu, \quad (3.9)$$

induce the following variation of Ω :

$$\delta\Omega = -d \left[\frac{1}{2} d\Lambda_{\mu\nu} J^{\mu\nu} \right], \quad (3.10)$$

where

$$J^{\mu\nu} = 2x^{[\mu} p^{\nu]} - \frac{1}{(mc)^2} \varepsilon^{\mu\nu\rho\sigma} p_\rho w_\sigma. \quad (3.11)$$

If we use this (and $P_\mu = p_\mu$) to compute the Pauli-Lubanski pseudo-vector L we find that

$$L^\mu := \frac{1}{2} \varepsilon^{\mu\nu\rho\sigma} P_\nu J_{\rho\sigma} = w^\mu, \quad (3.12)$$

and hence that

$$L^2 = w^2 = (mc)^2 s^2. \quad (3.13)$$

3.1 The tachyonic spinning particle and its Galilean limit

Now we consider the tachyonic version of Souriau's relativistic spinning particle model obtained by setting $mc = ik$ for some real number k . This yields the phase-space constraints

$$p^2 = k^2, \quad w^2 = -(ks)^2, \quad p \cdot w = 0. \quad (3.14)$$

Now p is spacelike and w is timelike, but the Souriau 2-form is still closed, by an application of the above lemma but with a reversed identification of (u, v) with multiples of (p, w) .

It is again true that Ω is not invertible on the 12-dimensional space parametrized by the components of the three 4-vectors (x, p, w) but *is* invertible on the 10-dimensional submanifold determined by the w -dependent constraints. However, the surface that these constraints define within the Minkowski subspace of fixed (x, p) is now a hyperboloid rather than a sphere. To see this we may choose a frame for which $p \propto (0, \mathbf{n})$ for unit 3-vector \mathbf{n} ; then w is a timelike vector of fixed interval in the 3D Minkowski subspace orthogonal to \mathbf{n} and hence lies on a 2-dimensional hyperboloid.

To take the Galilean limit of the spinning tachyon, we first rescale x^0, p_0 and \mathbf{w} as follows

$$x^0 \rightarrow \lambda x^0, \quad p_0 \rightarrow p_0/\lambda, \quad \mathbf{w} \rightarrow \mathbf{w}/\lambda, \quad (3.15)$$

where λ is positive, and then we take $\lambda \rightarrow \infty$. As $x^0 = ct$ and $p_0 = E/c$ (for dimensionless Minkowski metric) this is equivalent to the $c \rightarrow \infty$ limit but with an additional specification of how to take this limit for the components of w . One finds that the constraints (3.14) reduce to

$$|\mathbf{p}|^2 = k^2, \quad w_0^2 = (ks)^2, \quad p_0 w_0 = \mathbf{p} \cdot \mathbf{w}. \quad (3.16)$$

Assuming, for simplicity, that both w_0 and ks are positive, the second of these constraints tells us that $w_0 = ks$. The third constraint can be solved for the component of \mathbf{w} parallel to \mathbf{p} , but this leaves two components of \mathbf{w} undetermined. This is as expected because the w -dependent constraints initially restricted w to a 2-dimensional hyperboloid. However, when we perform the rescaling (3.15) in the action, and take the $\lambda \rightarrow \infty$ limit, these unrestricted variables drop out. If we use $w_0 = ks$ to eliminate w_0 , the Souriau 2-form reduces to the 2-form Ω_{WZ} of (2.2), and the net result is that we recover the action (2.1) for the Galilean massless particle of colour k and spin s .

A peculiar feature of this limit is that the physical phase space is only 6-dimensional in the limit whereas it was 8-dimensional initially. We suspect that this was the source of Souriau's reservations about the non-relativistic limit of a spinning tachyon.

4 Discussion

We have shown that the massless Galilean particle of colour k is a non-relativistic limit of a tachyon of mass $m = ik/c$. Although tachyons are usually considered to be unphysical, there are *unitary* irreducible tachyonic representations of the Poincaré group [9] and the possibility that these may have some physical realization has been explored in many papers; see e.g. [10] for a recent review with references to the literature. Consequently, one cannot conclude from its tachyonic origin that the massless Galilean particle is intrinsically unphysical.

However, this conclusion changes when we consider the supersymmetric extension of massless Galilean systems because there are no unitary irreducible tachyonic representations of the super-Poincaré group. The tachyonic superparticle is intrinsically non-unitary, and we should therefore expect the same of any attempt at a supersymmetrization of the massless Galilean particle. This argument does not apply to the Galilean superstring, for which unitarity simply requires the same bound (2.11) on the total momentum that is already implied by the classical phase-space constraints [5]. Inspection of the details shows that this is due to the intrinsically “stringy” topological charge in the super-Galilean algebra of Noether charges.

As a final comment, inspired by the idea expounded in [11] of a “duality” relating the Galilean to the Carroll limit [12], we observe that the status of a tachyon in the Galilean

limit is analogous to that of a bradyon in the Carroll limit [12]. In the former case, the superluminal velocity is forced to go to infinity. In the latter case the subluminal velocity must go to zero. In both cases, the mass (real or imaginary) corresponds to a property of the resulting system.

Note added:

Another massless Galilean system, in the Souriau sense, is provided by the particle with Lagrangian

$$L = \frac{\mu}{2} |\ddot{\mathbf{x}}|^2. \quad (4.1)$$

Because this Lagrangian is *strictly* Galilean invariant (its variation is *not* a total time derivative) there is no central charge in the algebra of Galilei Noether charges [13]; this also follows from dimensional analysis because the only parameter, μ , has dimensions of mass \times time-squared rather than mass. The corresponding phase-space Lagrangian is

$$L = \mathbf{p} \cdot \dot{\mathbf{x}} + \mathbf{q} \cdot \dot{\mathbf{y}} - H, \quad H = \mathbf{p} \cdot \mathbf{y} + \frac{1}{2\mu} |\mathbf{q}|^2, \quad (4.2)$$

where we use a rescaled version of the phase-space coordinates of [13]. By taking the $\mu \rightarrow \infty$ limit we get the phase-space Lagrangian

$$L_{SZ} = \mathbf{p} \cdot (\dot{\mathbf{x}} - \mathbf{y}) + \mathbf{q} \cdot \dot{\mathbf{y}}, \quad (4.3)$$

which was the basis for a dynamical alternative to dark energy proposed by Stichel and Zakrzewski [14]. These authors also considered a relativistic analog, which they interpreted as a tachyon. In order to elucidate the relation of this result to the results reported here, we present a brief analysis of the Stichel-Zakrzewski Lagrangian.

The equations of motion for \mathbf{y} and \mathbf{p} are jointly equivalent to

$$\mathbf{y} = \dot{\mathbf{x}}, \quad \mathbf{p} = \dot{\mathbf{q}}, \quad (4.4)$$

so we may consistently eliminate these variables to get an equivalent Lagrangian for \mathbf{x} and \mathbf{q} alone. In terms of the linear combinations

$$\mathbf{z}_{\pm} = \left(\mathbf{x} \mp \frac{1}{2m} \mathbf{q} \right), \quad (4.5)$$

where m is an arbitrary non-zero constant mass parameter, this equivalent Lagrangian is

$$L_{NR} = \frac{m}{2} [|\dot{\mathbf{z}}_+|^2 - |\dot{\mathbf{z}}_-|^2] + \frac{d}{dt} (\dots). \quad (4.6)$$

Each term is separately Galilean invariant, with Noether charges $\{\mathbf{P}_{\pm}, \mathbf{G}_{\pm}, \mathbf{J}_{\pm}\}$ and central charges $\pm m$. The linear combinations

$$\mathbf{P} = \mathbf{P}_+ + \mathbf{P}_-, \quad \mathbf{G} = \mathbf{G}_+ + \mathbf{G}_-, \quad \mathbf{J} = \mathbf{J}_+ + \mathbf{J}_-, \quad (4.7)$$

span a Galilei algebra with zero central charge because the total central charge is $m - m = 0$. So we indeed have a massless Galilean system, but at the cost of a non-unitary quantum theory.

The relativistic analog of L_{SZ} considered in [14] was presented as a set of equations to be satisfied by phase-space variables that were assumed to be functions of an arbitrary worldline time parameter, although the constraint generating time reparametrizations was not given. We can proceed more systematically now that we have established the equivalence of L_{SZ} to L_{NR} ; the latter is obviously the Galilean limit of the relativistic mechanics model with Lorentz invariant Lagrangian

$$L_{\text{Rel}} = -mc^2 \left[\sqrt{1 - |\mathbf{u}_+|^2} - \sqrt{1 - |\mathbf{u}_-|^2} \right], \quad (\mathbf{u}_\pm = \dot{\mathbf{z}}_\pm/c). \quad (4.8)$$

Each term is separately Lorentz invariant (although not manifestly so because the transformations are non-linear) and the $c \rightarrow \infty$ limit yields L_{NR} directly because the rest-mass energy cancels between the two terms. We also have two sets of Lorentz generators, in particular two conserved 4-momenta P_\pm and the Lorentz algebra with the Galilean limit is found by taking the sum. In particular, the total 4-momentum $P = P_+ + P_-$ is the combination relevant to the Galilean limit, and

$$P^2 = 2m^2c^2 [\gamma_+\gamma_- (1 - \mathbf{u}_+ \cdot \mathbf{u}_-) - 1] \geq 0, \quad \left(\gamma_\pm = \frac{1}{\sqrt{1 - |\mathbf{u}_\pm|^2}} \right). \quad (4.9)$$

It follows that P is spacelike unless $\mathbf{u}_+ = \mathbf{u}_-$, in which case it is null. This is the tachyonic behaviour found in [14] although we would choose to interpret the model as a two-particle system rather than a tachyon. In any case, the relative minus sign between the two terms of the Lagrangian L_{Rel} gives us information that is not obtainable from the equations of motion alone, and it tells us that the quantum theory is not unitary, as was to be expected from its Galilean limit.

Acknowledgements

We thank Peter Horvathy and Peter Stichel for bringing aspects of their work to our attention. LM is grateful for the hospitality of the Department of Applied Mathematics and Theoretical Physics at the University of Cambridge and the Departament de Física Quàntica i Astrofísica at the University of Barcelona. CB is partially supported by the Generalitat de Catalunya through project 2014 SGR 267 and by the Spanish government (MINECO/FEDER) under project CICYT DPI2015-69286-C3-2-R. JG has been supported in part by FPA2013-46570-C2-1-P, 2014-SGR-104 (Generalitat de Catalunya) and Consolider CPAN and by the Spanish government (MINECO/FEDER) under project MDM-2014-0369 of ICCUB (Unidad de Excelencia María de Maeztu). PKT acknowledges support from the UK Science and Technology Facilities Council (grant ST/L000385/1).

References

- [1] J.-M. Souriau, “Structure des systèmes dynamiques”, Dunod (1970, ©1969); “Structure of Dynamical Systems: A Symplectic View of Physics”, translated by C.H. Cushman-de Vries (R.H. Cushman and G.M. Tuynman, Translation Editors), Birkhäuser, 1997.
- [2] C. Duval, Z. Horvath and P. Horvathy, “Geometrical spinoptics and the optical Hall effect,” *J. Geom. Phys.* **57** (2007) 925 [math-ph/0509031].
- [3] C. Duval, “Polarized Spinoptics and Symplectic Physics,” arXiv:1312.4486 [math-ph].
- [4] C. Batlle, J. Gomis and D. Not, “Extended Galilean symmetries of non-relativistic strings,” *JHEP* **1702** (2017) 049 [arXiv:1611.00026 [hep-th]].
- [5] J. Gomis and P. K. Townsend, “The Galilean Superstring,” *JHEP* **1702** (2017) 105 [arXiv:1612.02759 [hep-th]].
- [6] J. F. Schonfeld, “A Mass Term for Three-Dimensional Gauge Fields,” *Nucl. Phys. B* **185** (1981) 157.
- [7] K. P. Tod, “Some Symplectic Forms Arising in Twistor Theory,” *Rept. Math. Phys.* **11** (1977) 339.
- [8] S. Fedoruk and J. Lukierski, “Massive twistor particle with spin generated by SouriauWessZumino term and its quantization,” *Phys. Lett. B* **733** (2014) 309 [arXiv:1403.4127 [hep-th]].
- [9] E. P. Wigner, “On Unitary Representations of the Inhomogeneous Lorentz Group,” *Annals Math.* **40** (1939) 149 [*Nucl. Phys. Proc. Suppl.* **6** (1989) 9].
- [10] V. F. Perepelitsa, “Looking for a theory of faster-than-light particles,” arXiv:1407.3245 [physics.gen-ph].
- [11] C. Duval, G. W. Gibbons, P. A. Horvathy and P. M. Zhang, “Carroll versus Newton and Galilei: two dual non-Einsteinian concepts of time,” *Class. Quant. Grav.* **31** (2014) 085016 [arXiv:1402.0657 [gr-qc]].
- [12] J.M. Lévy-Leblond, “Une nouvelle limite non-relativiste du group de Poincaré”, *Ann. Inst. H. Poincaré* **3** (1965) 1.
- [13] J. P. Gauntlett, J. Gomis and P. K. Townsend, “Particle Actions as Wess-Zumino Terms for Space-time (Super)symmetry Groups,” *Phys. Lett. B* **249** (1990) 255.
- [14] P. C. Stichel and W. J. Zakrzewski, “Can cosmic acceleration be caused by exotic massless particles?,” *Phys. Rev. D* **80** (2009) 083513 [arXiv:0904.1375 [astro-ph.CO]].